## **Cross-verification and validation of tokamak plasma evolution models**

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### **Summary**

- Validate Te/Ti predictors using state-of-theart but general settings
  - Run on ~hundreds of cases automatically
  - Compare multiple independent implementations (TRANSP and ASTRA)
  - Compare against empirical (linear regressed) models to contextualize error
- Find no significant statistical difference in Te/Ti predictions between TRANSP, ASTRA, and empirical model





- Inputs:
  - EFIT01 (no kinetic constraint) q
  - ZIPFIT ne, rotation, Zeff profiles
  - Te and Ti boundary at  $\rho = 0.8$
- 900ms simulation
- TGLF SAT2, same settings except nky

	TRANSP	ASTRA
Fast ions	NUBEAM	RABBIT
Equilibrium	(input directly)	SPIDER
Ion heat	+viscosity +cold-neutral CX	
Neoclassical diffusion	Modified Chang-Hinton	Angioni-Sauter (e) Galeev-Sagdeev (i)
TGLF nky	12	19



## Verification: ASTRA and TRANSP yield similar results

shot 191577, time=2.70s



#### • Exclude

- wave-heating
- 3D field perturbations
- Non-D2 gas
- Rampup and rampdown
- Shots before year 2010
- OMFIT modules
  - ASTRA: compiled + debugged + user-interface for GA (Iris) cluster
  - AGGregate: automatically mass prepare + launch TRANSP/ASTRA jobs





## ASTRA and TRANSP converge in ~half of cases, runs take ~hrs (wall-clock) ASTRA less robust, TRANSP higher runtime





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## **Metrics and figures of merit**

Metrics to consider: Te, Ti, and W<sub>MHD</sub>

 $W_{MHD} = \int (p_{thermal} + p_{fast ions}) dV$ 

#### **ITER standard figures of merit used to measure accuracy**





#### **Baselines for comparison:** W<sub>MHD</sub> from H89/98 (w/ linear regression) Te and Ti from profile consistency w/ linear regression

"Nondimensionalization"1

 $\widehat{W}_{MHD} = P_{tot}\tau_{H\{89,98\}}$  $P_{tot} = P_{NBI} + 0.55MW$ 



1988 Profile Consistency: Global and nonlinear transport <sup>3</sup>Coppi

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## TRANSP and ASTRA qualitatively capture time-dependent changes





## Full database: TRANSP/ASTRA within $\sim 5\%$ Empirical Te/Ti also within $\sim 5\%$ , but $W_{MHD} > \sim 10\%$ worse







## Conclusions and next steps: as we know, codes just one component to predict

- Developers are aware of models' limitations: primarily used not predictively but for
  - Qualitative scaling
  - Physics understanding
  - Extrapolation to unexplored regimes
- BUT similar transport prediction workflows still used to plan
  - Reactors
  - Scenarios
  - Machine upgrades
- In practice, code outputs are combined with experience + empirical scalings
- Use machine learning to try a task humans have always done:
  - More rigorously understand where and when to trust codes vs empirical data
  - Maintain extrapolability to new regimes with power of empirical models
- Start by predicting difference from code to experiment value (w/ database we made)



- Expt title: "Effects of upstream power and heat flux width on the SAS-VW heat flux profile; influence of radiative and neutral heating"
  - Detachment studies
- Al Hyatt (shot log): "Very strange behavior. betan and density and li all seem to oscillate at a few hertz until the plasma density reaches about 4-5+13. Strike is almost perfect, maybe a little (~1 mm) too far out."



## **Detailed heat source comparison**





## **Detailed TGLF settings**

- sat\_rule: 2
- use\_bper but not bpar
- kygrid\_model: 1
- wdia\_transp: 1
- xnu\_model: 3
- alpha\_quench: 0
- n\_species: 3 (electrons, ions, impurity)
- n\_modes: 3
- ibranch=-1
- etg factor: 1.25
- gaussian width: 1.65
- growth rate search for max width from 0.3 to 21
- units: cgyro

